




The Marshall and Olkin-G and Gamma-G Family of Distributions: Properties and Applications

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Abstract:

- The article introduces a new family by combining the Marshall and Olkin-G and Gamma-G classes. It has only two extra shape parameters and can be a better model than other existing classes of distributions. Simulations are performed to verify the consistency of the estimators. Its flexibility is shown by means of two real data sets.

Keywords:

- *applications; distribution family; mathematical properties; simulations.*

AMS Subject Classification:

- 60E05, 62E10, 62E15.

1. INTRODUCTION

The advances in the field of Data Science requires the search for new families of distributions that adequately model real data has been increasing steadily in the last years. The construction of different generation methods and even generators of families has made it difficult to compare new proposals. In the midst of a huge set of existing families in the literature of new distributions, to find a proposal that is in fact an excellent competitor when compared to other existing ones, in terms of adjustment to real data sets and also that does not present estimation problems, is a major challenge.

The classes of distributions in the early 1980s were based on the simple idea of adding parameters to a baseline distribution. The mechanism by adding shape parameters to a baseline distribution has proved to be useful to make the generated distributions more flexible especially for studying tail properties than existing ones and for improving their goodness-of-fit statistics to real data. Many special distributions in these families are discussed by Tahir and Nadarajah (2015) [1].

The addition of parameters in the construction of new distributions/families was improved by the inclusion of mathematical functions known in the literature, such as beta and gamma functions, for example, which produce new generators with more flexible properties than their baselines. Two well-known examples are the beta-G (Eugene *et al.*, 2002) [2] and gamma-G (Zografos and Balakrishnan, 2009) [3] generators.

However, the inclusion of such functions for generating new families brought, in some cases, problems for parameter estimation. So, despite the fit being more suitable for some types of data and, therefore, having a superior performance when compared to other generators, the estimation process can often be a problem.

In this context, this work presents a new family obtained by composing a very competitive class in the literature with another class that has the gamma function in its structure.

Let $G(x)$ be the cumulative distribution function (CDF) of a baseline distribution and $g(x) = dG(x)/dx$ be the corresponding probability density function (PDF) depending on a parameter vector η . A generalized family is presented with two extra shape parameters by transforming the CDF $G(x)$ according to two sequential important classes. These classes, called Marshall and Olkin-G and Gamma-G, are important for modeling data in several areas, and they are reviewed below.

The CDF of the Marshall and Olkin's (1997) [4] (MO-G) class (for $\theta > 0$) is

$$(1.1) \quad F_{\text{MO-G}}(x) = \frac{G(x)}{\theta + (1 - \theta)G(x)} = \frac{G(x)}{1 - (1 - \theta)[1 - G(x)]}, \quad x \in \mathbb{R}.$$

The density function corresponding to (1.1) has the form

$$(1.2) \quad f_{\text{MO-G}}(x) = \frac{\theta g(x)}{[\theta + (1 - \theta)G(x)]^2}.$$

For $\theta = 1$, $f_{\text{MO-G}}(x)$ is equal to $g(x)$. Equation (1.2) represents the PDF of the minimum of n iid random variables having density $g(x)$, say T_1, \dots, T_N , where N has a geometric distribution with probability parameters θ and θ^{-1} if $0 < \theta < 1$ and $\theta > 1$, respectively.

Tahir and Nadarajah (2015, Table 2) [1] presented thirty distributions belonging to this family. It is easily generated from the baseline quantile function (QF) by $Q_{\text{MO-G}}(u) = Q_G(\theta u / [\theta u + 1 - u])$ for $u \in (0, 1)$.

Marshall and Olkin considered the exponential and Weibull distributions for the baseline G and derived some structural properties of the generated distributions. If G is an exponential distribution, the special case refers to a two-parameter competitive model to the Weibull and gamma distributions.

The CDF of the gamma-G (Γ -G) class (Zografos and Balakrishnan, 2009) [3] is

$$(1.3) \quad F_{\Gamma\text{-G}}(x) = \gamma_1(a, -\log[1 - G(x)]), \quad x \in \mathbb{R},$$

where $a > 0$ is an extra shape parameter, $\gamma_1(a, z) = \gamma(a, z) / \Gamma(a)$ is the incomplete gamma function ratio, and $\gamma(a, z) = \int_0^z t^{a-1} e^{-t} dt$.

Then, the PDF of the Γ -G class can be expressed as

$$(1.4) \quad f_{\Gamma\text{-G}}(x) = \frac{1}{\Gamma(a)} \{-\log[1 - G(x)]\}^{a-1} g(x).$$

Each new Γ -G distribution follows from a given baseline G . For $a = 1$, the Γ -G class reduces to G . If Z is a gamma random variable with unit scale and shape $a > 0$, then $W = Q_G(1 - e^{-Z})$ has density (1.4). So, the Γ -G distribution is easily generated from the gamma distribution and the QF of G .

The remaining of the paper is addressed as follows. Section 2 introduces the *Marshall and Olkin-Gamma-G* (MOGa-G) family, and provides some special models. The maximum likelihood estimates (MLEs) of its parameters is addressed in Section 3. Some simulations are performed in Section 4 to estimate the biases of the MLEs. Two empirical applications illustrate the potentiality of the proposed family in Section 5. A variety of theoretical properties are obtained in Section 6. Some conclusions remarks are offered in Section 7.

2. THE NEW FAMILY

By combining Equations (1.1) and (1.3), the CDF of the random variable $X \sim \text{MOGa-G}$ representing the new family has the form

$$(2.1) \quad F_X(x) = \frac{\gamma_1(a, -\log[1 - G(x)])}{\theta + (1 - \theta)\gamma_1(a, -\log[1 - G(x)])}, \quad x \in \mathbb{R}.$$

By differentiating (2.1), the PDF of X follows as

$$(2.2) \quad f_X(x) = \frac{\theta \{-\log[1 - G(x)]\}^{a-1} g(x)}{\Gamma(a) \{\theta + (1 - \theta)\gamma_1(a, -\log[1 - G(x)])\}^2}.$$

The density (2.2) can be interpreted from a sequence of N iid random variables, say Z_1, \dots, Z_N , each one having a gamma density with unit scale and shape $a > 0$, assuming that N (is not fixed) has a geometric distribution with probabilities θ and θ^{-1} for $0 < \theta < 1$ and $\theta > 1$, respectively. By transforming the Z_i 's via the baseline QF by $W_i = Q_G(1 - e^{-Z_i})$ (for $i = 1, \dots, N$), Equation (1.2) defines the PDF of the minimum W_1, \dots, W_n . The proposed family from the double composition of the two classes absorbs the impacts of their different flexibilities on real applications.

Table 1 provides some special cases of (2.2), where $\Phi(x)$ and $\phi(x)$ are the CDF and PDF of the standard normal distribution. The density and hazard ($h(x) = f(x)/[1 - F(x)]$) functions of the MOGa-Weibull (MOGa-W) model are displayed in Figure 1, which provide more flexibility to both functions for both classes applied separately to the Weibull model.

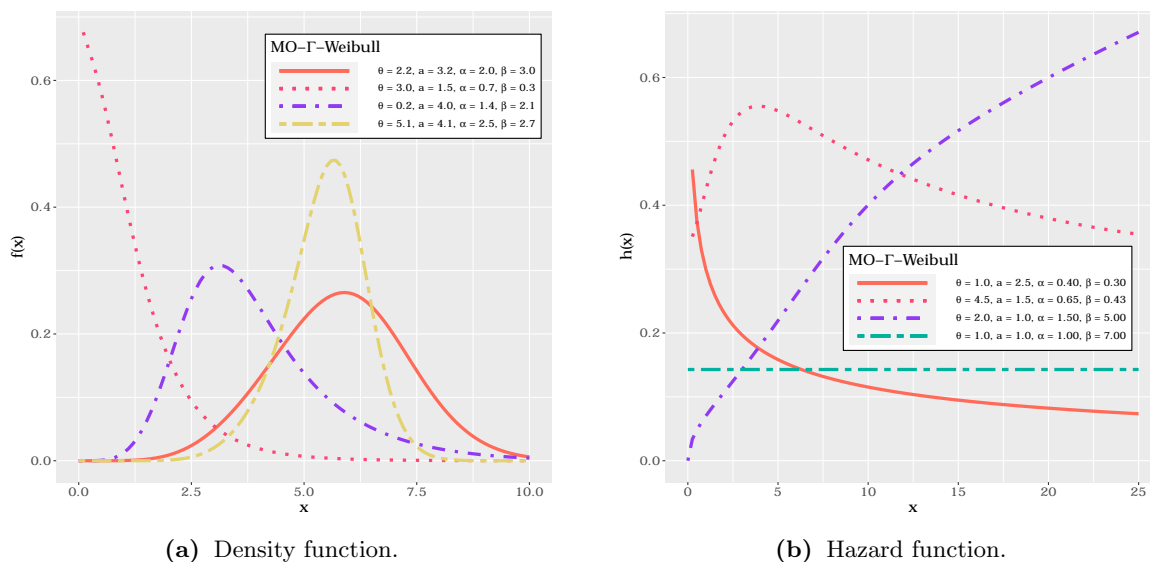


Figure 1: The density and hazard functions of the MOGa-W model.

The CDF (2.1) can be easily inverted to calculate the QF of the MOGa-G distribution, say $x = Q_X(u) = F_X^{-1}(u)$ (for $u \in (0, 1)$), in terms of the baseline QF $Q_G(\cdot)$. The inverse of $F_X(x) = u$, where u is a uniform number in $(0, 1)$, follows by combining the inverses of Equations (1.1) and (2.1). So, $F_X(x) = u$ gives $z = z(u) = \theta u/[1 - (1 - \theta)u]$ and $\gamma_1(a, -\log[1 - G(x)]) = z(u)$. Hence, the QF of X can be expressed as

$$x = Q_G(v(u)),$$

where

$$v(u) = 1 - \exp[-\gamma_1^{-1}(a, z(u))],$$

and $\gamma_1^{-1}(a, w) = Q^{-1}(a, 1 - w)$ is the inverse function of $\gamma_1(a, w)$. Some formulae for $Q^{-1}(a, 1 - w)$ are given in <http://functions.wolfram.com/GammaBetaErf/InverseGammaRegularized/>.

Table 1: Special Distributions in the MOGa-G family.

Distribution	Baseline CDF	Generated PDF
Normal	$G(x) = \Phi(x)$	$f_X(x) = \frac{\theta \{-\log[1-\Phi(x)]\}^{\alpha-1} \phi(x)}{\Gamma(a) \{\theta+(1-\theta)\gamma_1(a, -\log[1-\Phi(x)])\}^2}$
Logistic	$G(x) = \frac{1}{1+e^{-x}}$	$f_X(x) = \frac{\theta e^{-x} \{-\log[1-(1+e^{-x})^{-1}]\}^{\alpha-1}}{\Gamma(a) (1+e^{-x})^2 \{\theta+(1-\theta)\gamma_1(a, -\log[1-(1+e^{-x})^{-1}])\}^2}$
Gumbel	$G(x) = 1 - \exp(-e^x)$	$f_X(x) = \frac{\theta \exp(ax - e^x)}{\Gamma(a) \{\theta+(1-\theta)\gamma_1(a, e^x)\}^2}$
Log-Normal	$G(x) = \Phi(\log x)$	$f_X(x) = \frac{\theta \phi(\log x) \{-\log[1-\Phi(\log x)]\}^{\alpha-1}}{\Gamma(a) x \{\theta+(1-\theta)\gamma_1(a, -\log[1-\Phi(\log x)])\}^2}$
Exponential	$G(x) = 1 - \exp(-\lambda x), \lambda > 0$	$f_X(x) = \frac{\theta \lambda^\alpha x^{\alpha-1}}{\Gamma(a) \{\theta+(1-\theta)\gamma_1(a, \lambda x)\}^2}$
Weibull	$G(x) = 1 - \exp(-(\lambda x)^\gamma), \lambda, \gamma > 0$	$f_X(x) = \frac{\theta \gamma \lambda^\alpha \gamma x^{\alpha \gamma - 1} \exp\{-(\lambda x)^\gamma\}}{\Gamma(a) \{\theta+(1-\theta)\gamma_1[a, (\lambda x)^\gamma]\}^2}$
Gamma	$G(x) = \gamma_1(\alpha, \beta x), \alpha, \beta > 0$	$f_X(x) = \frac{\theta \beta^\alpha x^{\alpha-1} e^{-\beta x} \{-\log[1-\gamma_1(\alpha, \beta x)]\}^{\alpha-1}}{\Gamma(a) \{\theta+(1-\theta)\gamma_1(a, -\log[1-\gamma_1(\alpha, \beta x)])\}^2}$
Pareto	$G(x) = 1 - \frac{1}{(1+x)^\nu}, \nu > 0$	$f_X(x) = \frac{\theta e^{-x} [\nu \log(1+x)]^{\alpha-1} g(x)}{\Gamma(a) (1+e^{-x})^2 \{\theta+(1-\theta)\gamma_1(a, \nu \log[1+x])\}^2}$
Dagum	$G(x) = [1 + (x/\beta)^{-\alpha}]^{-p}, \alpha, \beta, p > 0$	$f_X(x) = \frac{\theta \{-\log[1-[1+(x/\beta)^{-\alpha}]^{-p}]\}^{\alpha-1} g(x)}{\Gamma(a) \{\theta+(1-\theta)\gamma_1[a, -\log(1-((x/\beta)^{-\alpha}+1)^{-p})]\}^2}$

3. ESTIMATION

The MOGa-G family can be fitted to real data using the **AdequacyModel** package (Marinho *et al.*, 2019) [5] in the R software. This package does not require to define the log-likelihood function, and it computes the MLEs, their standard errors (SEs), and the formal statistics defined in Section 5. It is only necessary to provide the PDF and CDF of the distribution to be fitted to a data set.

For example, if x_i is one observation from (2.2) and $\boldsymbol{\eta}$ is a q -parameter vector specifying $G(\cdot)$ as the Weibull CDF, the log-likelihood function for $\boldsymbol{\theta}^\top = (a, \theta, \boldsymbol{\eta}^\top)$ from n observations can be expressed as

$$\begin{aligned}
 \ell(\boldsymbol{\theta}) = & n \log(\theta) + n \log(\gamma) + n a \gamma \log(\lambda) + (a\gamma - 1) \sum_{i=1}^n \log(x_i) - \lambda^\gamma \sum_{i=1}^n x_i^\gamma \\
 (3.1) \quad & - n \log[\Gamma(a)] - 2 \sum_{i=1}^n \log\{\theta + (1 - \theta)\gamma_1[a, (\lambda x_i)^\gamma]\}.
 \end{aligned}$$

Due to the impossibility of obtaining the MLEs in closed form, numerical methods to calculate the estimates that maximize $\ell(\cdot)$ are necessary. Several programming languages and statistical software provides functions and routines that make it easy to obtain numerical estimates by various interactive methods. In practice, these estimates are commonly found in this way, since the Newton and quasi-Newton methods produce satisfactory results under reasonable conditions of the object function, i.e., when they do not impose restrictions that disturb the convergence of the algorithms.

The **AdequacyModel** package of the programming language R is used to obtain the MLEs, see R Core Team (2020) [6]. This library, created and maintained by one of the authors of this paper, is widely cited by several works in statistics, and serves as a basis for other library implementations available on the Comprehensive R Archive Network (CRAN). By using the `goodness.fit` function, it is possible to provide an implementation of (2.2), and obtain $\ell(\cdot)$ by returning the MLEs, and some measures of adequacy of fit. Further details regarding this package can be obtained from Marinho *et al.* (2019) [5].

4. SIMULATIONS

Due to the probable absence of MLEs in closed-form for distributions belonging to the MOGa-G family, it is necessary to examine the precision of the estimates calculated numerically.

In order to do that, the biases of the estimators of the parameters of the MOGa-Dagum($\theta, a, \alpha, \beta, p$) and MOGa-Weibull($\theta, a, \lambda, \gamma$) distributions are determined, where $G \sim \text{Dagum}(\alpha, \beta, p)$ and $\text{Weibull}(\gamma, \lambda)$ are the baseline models, respectively. All parameters are taken equal to one for different sample sizes reported in Tables 2 and 3.

The numbers in Tables 2 and 3 indicate that the estimation method behaves well when the sample size increases. This is theoretically expected. However, in practice, difficulties can be faced to other families due to the flatness of the log-likelihood function.

All Monte Carlo simulations can be reproduced using the script in <https://github.com/prdm0/MOGG>. The simulations are parallelized and able to use all threads available by a multicore processor, thus making them more computationally efficient, and consequently requiring less time to complete.

The simulations are performed on a computer with an Intel(R) Core(TM) i5-9500 CPU processor with 6 threads working at a maximum frequency of 3.00GHz, requiring, on this hardware, a time of 15.4828 hours to perform all simulations, 7.7414 hours for the MOGa-Dagum($\theta, a, \alpha, \beta, p$) distribution, and 4.9688 hours for the MOGa-Weibull($\theta, a, \lambda, \gamma$) distribution. Tables 2 and 3 reveal that the average biases of the MLEs could be very small for $n > 2,000$.

To generate observations from the random variable X with density f , the well-known Acceptance-Rejection Algorithm for continuous random variables is very useful when the QF involves complex functions that can lead to some numerical inaccuracies. For doing this, another random variable Y is chosen such that it can generate observations from a PDF h with the same support as f . Then, the acceptance and rejection algorithm is defined by the following steps:

1. Generate an outcome y from Y ;
2. Generate an observation u from a random variable $U \sim \mathcal{U}(0, 1)$;
3. If $u < \frac{f(y)}{cg(y)}$, where c is a real constant, accept $x = y$; otherwise reject y as an outcome from X and return to 1.

The constant c must be chosen in such a way that $\frac{f(y)}{cg(y)} \leq 1$. Thus, to minimize the computational cost of generating observations from X through the generated observations from Y , c is chosen as the lowest possible value to maximize the likelihood of acceptance. Further details of this method can be found in Rizzo (2019) [7].

Table 2: Average biases of the MLEs of the MOGGa-Dagum($\theta, a, \alpha, \beta, p$) distribution calculated by the BFGS method from simulations.

n	$B(\hat{\theta})$	$B(\hat{a})$	$B(\hat{\alpha})$	$B(\hat{\beta})$	$B(\hat{p})$	Time (mins)
10	0.2213	2.1944	2.6971	1.5803	1.3190	0.6960
20	0.4240	2.4793	1.5591	1.8083	0.7414	0.9819
60	0.7458	2.2661	0.5598	1.8495	0.2812	1.9417
100	0.6194	1.9438	0.3312	1.6142	0.2935	2.7208
200	0.3950	1.4262	0.1856	1.1556	0.3611	4.4534
400	0.2077	0.9599	0.1082	0.6157	0.4076	7.4698
600	0.1200	0.7213	0.0767	0.4024	0.3572	9.4975
1000	0.0629	0.4791	0.0503	0.2123	0.2584	12.4221
2000	0.0362	0.2958	0.0298	0.1145	0.1878	20.7251
5000	-0.0040	0.1325	0.0159	0.0144	0.0167	28.3380
10000	-0.0133	0.0815	0.0096	0.0081	0.0039	50.9298
20000	-0.0111	0.0349	0.0037	0.0006	-0.0109	68.6320
30000	-0.0036	0.0191	0.0006	-0.0041	-0.0034	97.3046
50000	-0.0057	0.0129	0.0016	0.0015	-0.0026	158.3737

Table 3: Average biases of the MLEs of the MOGGa-Weibull($\theta, a, \lambda, \gamma$) distribution calculated by the BFGS method from simulations.

n	$B(\hat{\theta})$	$B(\hat{a})$	$B(\hat{\lambda})$	$B(\hat{\gamma})$	Time (mins)
10	0.0818	0.1362	4.9274	1.2407	0.6716
20	0.3404	-0.0177	3.4160	1.4117	0.8077
60	0.7037	-0.0677	1.8806	1.3385	1.1773
100	0.6698	-0.0535	1.3684	1.1796	1.2643
200	0.5371	-0.0299	0.8265	0.9110	1.7886
400	0.3371	-0.0047	0.4205	0.5967	2.8386
600	0.2457	0.0076	0.2685	0.4306	3.6867
1000	0.1476	0.0093	0.1553	0.2818	5.0944
2000	0.0731	0.0035	0.0758	0.1530	8.8577
5000	0.0264	0.0007	0.0283	0.0618	15.8586
10000	0.0128	-0.0007	0.0142	0.0318	29.8629
20000	0.0053	-0.0012	0.0071	0.0160	48.7417
30000	0.0023	-0.0014	0.0053	0.0119	64.7387
50000	0.0023	-0.0004	0.0028	0.0063	112.7422

5. APPLICATIONS

Two applications compare the MOGGa-Weibull (MOGGa-W for short) model with seven extended Weibull distributions: the beta-Weibull (β -W) (Famoye *et al.*, 2005) [8], Kumaraswamy Weibull (Kw-W) (Cordeiro *et al.*, 2010) [9], Marshall-Olkin Weibull (MO-W) (Ahmed *et al.*, 2017) [10], Marshall-Olkin Extended Weibull (MOE-W) (Cordeiro *et al.*,

2019) [11], exponentiated Weibull (exp-W) (Mudholkar and Srivastava, 1993) [12], gamma Weibull (Γ -W) (Cordeiro *et al.*, 2016) [13], and exponentiated generalized Weibull (EG-W) (Oguntunde *et al.*, 2015) [14] (with $a = 1$). Some of these distributions are widely used in practice.

The log-likelihood for $\boldsymbol{\theta}$ from the MOGa-W distribution from one observation can be expressed as

$$(5.1) \quad \begin{aligned} \ell(\boldsymbol{\theta}) = & \log(\theta) + \log(\gamma) + (a\gamma)\log(\lambda) + (a\gamma - 1)\log(x) - (\gamma x)^\gamma - \log[\Gamma(a)] \\ & - 2\log\{\theta + (1 - \theta)\gamma_1[a, (\lambda x)^\gamma]\}, \end{aligned}$$

where $\boldsymbol{\theta} = (a, \theta, \lambda, \gamma)^\top$. The components of the score function are

$$U_a(\boldsymbol{\theta}) = \gamma \log(\lambda) + \gamma \log(x) - \psi^{(0)}(a) - \frac{2\{(1 - \theta)A - (1 - \theta)\psi^{(0)}(a)\gamma_1[a, (\lambda x)^\gamma]\}}{\theta \Gamma(a) + (1 - \theta)\gamma_1[a, (\lambda x)^\gamma]},$$

$$U_\theta(\boldsymbol{\theta}) = \frac{1}{\theta} - \frac{2\{\Gamma(a) - \gamma_1[a, (\lambda x)^\gamma]\}}{\theta \Gamma(a) + (1 - \theta)\gamma_1[a, (\lambda x)^\gamma]},$$

$$U_\lambda(\boldsymbol{\theta}) = \frac{\gamma}{\lambda}[a - (\lambda x)^\gamma] + \frac{2\gamma \lambda^{-1}(\lambda x)^{a\gamma}(1 - \theta) \exp\{-(\lambda x)^\gamma\}}{\theta \Gamma(a) + (1 - \theta)\gamma_1[a, (\lambda x)^\gamma]}$$

and

$$U_\gamma(\boldsymbol{\theta}) = \frac{1}{\gamma} + a \log(\lambda) + a \log(x) - (\lambda x)^\gamma \log(\lambda x) + \frac{2(1 - \theta)(\lambda x)^{\gamma a} \log(\lambda x) \exp\{-(\lambda x)^\gamma\}}{\theta \Gamma(a) + (1 - \theta)\gamma_1[a, (\lambda x)^\gamma]},$$

where $\psi^{(n)}(x)$ is the n -th derivative of the digamma function,

$$A = \log[(\lambda x)^\gamma] \gamma_1[a, (\lambda x)^\gamma] + G_{2,3}^{3,0} \left((\lambda x)^\gamma \left| \begin{matrix} 1, 1 \\ 0, 0, a \end{matrix} \right. \right),$$

and $G_{p,q}^{m,n} \left(z \left| \begin{matrix} a_1, \dots, a_p \\ b_1, \dots, b_q \end{matrix} \right. \right)$ is the Meijer G function.

The **AdequacyModel** package is used to fit the previous distributions to two real data sets. The SANN method, which is a variant of simulated annealing algorithm (Belisle, 1992) [15], is used here. The distributions are compared via the Anderson Darling (A^*) and Cramér von Mises (W^*) statistics reported in the `goodness.fit` function.

For the first case, the **betareg** package is applied to a modification of the ‘‘FoodExpenditure’’ data, which refer to the proportions of income spent on food for 38 households in a large US city (according to the package information). The household expenditures for food are given by

$$data = FoodExpenditure_{food} / \#(FoodExpenditure_{food}),$$

where $FoodExpenditure_{food}$ is the random variable corresponding to the household expenditures for food, and $\#(\cdot)$ indicates the number of observations on this variable. The observations for the first data set are given below:

0.4210000, 0.4382105, 0.5721316, 0.1955526, 0.2758158, 0.3565263, 0.6120000, 0.4730526, 0.3726579, 0.2322368, 0.3732632, 0.5158947, 0.3612632, 0.5563421, 0.4591053, 0.2533947, 0.3685526, 0.2410526, 0.4955526, 0.2010789, 0.3653158, 0.2544737, 0.5685263, 0.2859474, 0.7626316, 0.2863684, 0.4884474, 0.3060263, 0.4754474, 0.3826053, 0.5050526, 0.6820526, 0.7587632, 0.4176053, 0.3923684, 0.2513158, 0.6070000, 0.3881842.

Some descriptive statistics are reported in Table 4. The minimum value refers to a family of 3 people and income of 39,151, which does not represent the lowest family income for the current data, as expected, occupying only the fifth position among those with the lowest income. The maximum value corresponds to a family of 6 people with an income of 69,929, the second-largest number among the number of people per family in the group in question. Furthermore, we can note that the current data present positive asymmetry and negative kurtosis.

Table 4: Descriptive statistics for the food data.

Minimum	0.1956
1st Qu.	0.2913
Median	0.3903
Mean	0.4198
3rd Qu.	0.5027
Maximum	0.7626
Standard Deviation	0.1480
Skewness	0.5250
Kurtosis	-0.4440

In addition, the standard deviation is relatively low. Figure 2 displays the total time on test (TTT) plot for the first data set, which shows that the failure function is decreasing. So, the MOG_a-W distribution is appropriate to fit these data, since its hazard function presents this shape (see, Figure 1(b)).

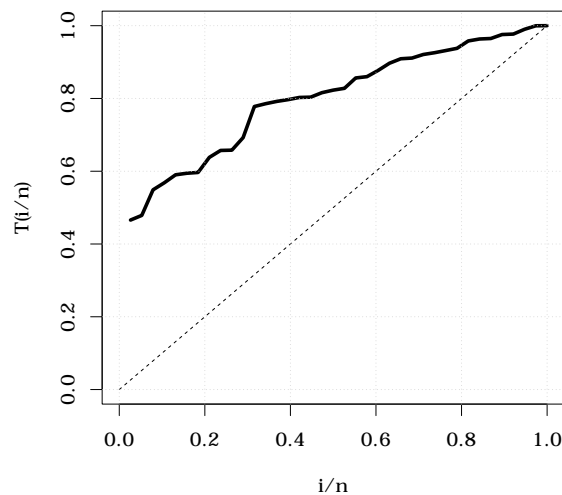


Figure 2: TTT plot for the food data.

The MLEs, their standard errors (SEs) (in parentheses), and the statistics W^* and A^* for the fitted models to the current data are listed in Table 5. The results indicate that the proposed model has better performance than the other seven fitted models.

Table 5: Estimation results for food data.

Model	\hat{a}	$\hat{\theta}$	$\hat{\lambda}$	$\hat{\gamma}$	W^*	A^*
MOGa-W($a, \theta, \lambda, \gamma$)	0.9261 (0.0262)	1.3796 (0.2238)	33.3230 (0.2853)	25.3988 (0.0825)	0.0339	0.2376
β -W($a, \theta, \lambda, \gamma$)	9.9288 (0.0290)	0.1700 (0.0204)	9.7594 (<0.0001)	1.5305 (<0.0001)	0.0435	0.2594
KW-W($a, \theta, \lambda, \gamma$)	0.0498 (0.0080)	99.9998 (16.2259)	1.0760 (0.0031)	23.4028 (0.0146)	1.3309	6.7426
MOE-W($a, \theta, \lambda, \gamma$)	0.1366 (0.1599)	2.0204 (<0.0001)	62.7220 (<0.0001)	4.2956 (0.7365)	0.0354	0.2579
EGW(a, b, λ, γ)	5.6189 (0.0028)	6.1833 (0.0009)	1.2870 (0.1159)	1.3798 (0.1480)	0.0371	0.2518
MO-W(a, λ, γ)	0.1592 (0.0717)	— (—)	1.5860 (0.1335)	4.2671 (0.1665)	0.0345	0.2573
exp-W(a, λ, γ)	6.1102 (0.4222)	— (—)	4.4680 (0.3175)	1.3858 (0.1677)	0.0372	0.2523
γ -W(a, λ, γ)	5.7515 (0.0015)	— (—)	10.0000 (0.0001)	1.2087 (0.0082)	0.0879	0.6599

A data set collected in a pilot study about hypertension in the Dominican Republic in 1997 refers to the second application. The observations are the systolic blood pressure of persons who came to medical clinics in several villages for a variety of complaints. The observations for the data set in question are:

150, 120, 120, 180, 138, 115, 130, 150, 200, 120, 190, 90, 130, 120, 200, 140, 110, 134, 160, 140, 105, 126, 129, 120, 100, 130, 118, 144, 180, 138, 110, 140, 120, 118, 110, 110, 130, 140, 130, 165, 180, 130, 140, 112, 130, 158, 112, 150, 140, 142, 110, 140, 130, 132, 140, 140, 122, 128, 90, 118, 120, 110, 122, 200, 110, 140, 150, 120, 150, 120, 164, 122, 112, 130, 140, 102, 122, 130, 102, 130, 122, 200, 140, 180, 124, 110, 124, 90, 120, 159, 142, 140, 118, 122, 108, 170, 120, 140, 100, 118, 110, 114, 150, 160, 140, 190, 118, 120, 150, 120, 200, 150, 168, 110, 142, 150, 160, 142, 160, 150, 110, 128, 122, 150, 140, 122, 120, 130, 100, 130, 150, 130, 100, 120, 105, 100, 150, 196, 130, 110, 140, 122, 110, 164, 120, 120, 150, 160, 150, 135, 124, 110, 100, 95, 130, 120, 108, 118, 170, 105, 120, 95, 95, 120, 140, 142, 160, 110, 190, 180, 130, 130, 120, 204, 150, 150, 120, 122, 120, 130, 140, 148, 118, 126, 136, 140, 130, 102, 110, 110, 130, 126, 142, 140, 128, 130, 124, 162, 130, 130, 110, 80, 166, 140, 160, 160, 140, 98, 138, 120, 112, 112, 134, 140, 115, 140, 98, 115, 120, 80, 160, 126, 110, 130, 104, 236, 118, 120, 140, 120, 98, 164, 150, 110, 120, 130, 170, 180, 110, 120, 130, 118, 130, 190, 158, 90, 99, 210, 180, 140, 184, 105, 120, 150, 140, 130, 160, 118, 210, 100, 170, 150, 130, 170, 150, 120, 134, 90, 125, 170, 140, 150, 110, 105, 140, 120, 100, 124, 112, 160, 140, 118, 190, 110, 118, 160, 150, 124, 128, 150, 120, 125, 118, 132, 110, 143, 170, 98, 124, 180, 178, 110, 98, 159, 110, 140, 130, 122, 110, 98, 180, 90, 118, 165, 138, 138, 170, 106, 170, 140, 90, 118, 110, 102, 102, 180, 100, 110, 162, 140, 110, 98, 140, 140, 110, 170, 112, 90, 102, 106, 124, 110, 180, 138, 90, 150, 126, 110, 130, 150, 145, 140, 156, 110, 150, 160, 120, 140, 120, 110, 120, 140, 160, 160, 110, 150, 118, 110, 120, 120, 146, 124, 170, 124, 170, 159, 120, 120, 118, 152, 190.

Table 6 provides the descriptive statistics for the second data set. Considering that the systolic blood pressure represents the highest number presented in the pressure measuring equipment, the maximum (236) found in this table should represent an individual with serious heart problems. This is due to the fact that the normal systolic pressure is 120.

Table 6: Descriptive statistics for the clinic data.

Minimum	80
1st Qu.	118
Median	130
Mean	133
3rd Qu.	150
Maximum	236
Standard Deviation	25.7157
Skewness	0.7893
Kurtosis	0.5908

The minimum value (80) should be for an individual who probably suffers from low blood pressure. Also note that the data set has positive skewness. This indicates that few people have high pressure values and, in this case, the mode is 120, which represents a good value for systolic pressure.

Figure 3 provides the TTT plot for the second clinic data, which shows that the hazard function is decreasing, thus supporting the MOGGa-W model.

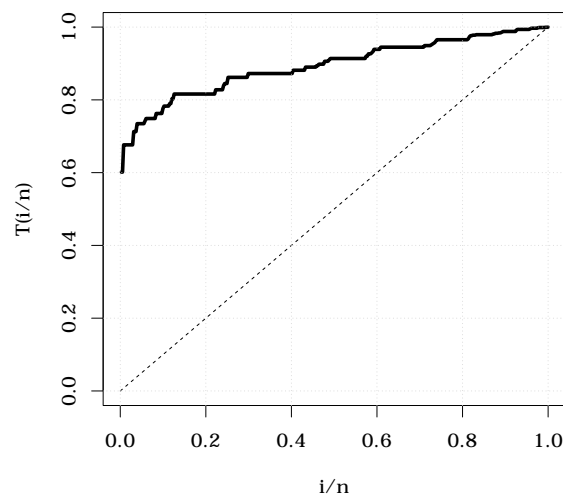


Figure 3: TTT plot for the clinic data.

The MLEs of the parameters, their SEs and the values of the adequacy measures for the fitted models to the clinic data are reported in Table 7. By comparing the measure values, the proposed distribution outperforms all other fitted models.

Table 7: Estimation results for clinic data.

Model	\hat{a}	$\hat{\theta}$	$\hat{\lambda}$	$\hat{\gamma}$	W^*	A^*
MOGa-W($a, \theta, \lambda, \gamma$)	9.6293 (0.0062)	3.6407 (0.1824)	6.2608 (0.0244)	12.8234 (0.0079)	0.5093	2.8076
β -W($a, \theta, \lambda, \gamma$)	31.0847 (0.0127)	47.1463 (<0.0001)	0.0169 (0.0001)	2.0540 (<0.0001)	0.7540	4.2794
KW-W($a, \theta, \lambda, \gamma$)	7363.2810 (0.0419)	0.0392 (0.0004)	1.4676 (<0.0001)	0.6146 (<0.0001)	0.5351	2.9617
MOE-W($a, \theta, \lambda, \gamma$)	101.1347 (46.7828)	0.4238 (0.1008)	0.0309 (0.0036)	1.7024 (0.1681)	1.1441	6.6446
EGW(a, b, λ, γ)	0.2351 (0.0025)	140.0000 (3.2785)	0.4576 (<0.0001)	0.7425 (<0.0001)	0.8925	5.3338
MO-W(a, λ, γ)	173.2139 (0.0001)	— (—)	0.0212 (0.0002)	1.6019 (0.0001)	1.4760	8.5870
exp-W(a, λ, γ)	69.0291 (0.0838)	— (—)	0.0240 (0.0002)	1.3455 (<0.0001)	0.8899	5.0941
Γ -W(a, λ, γ)	9.1122 (0.9633)	— (—)	0.0261 (0.0029)	1.7464 (0.0803)	0.6227	3.4882

6. MATHEMATICAL PROPERTIES

Here, some mathematical properties for the MOGa-G family are presented based on a linear representation for its density function in terms of “exponentiated-G” (exp-G) densities.

6.1. Linear Representation

For an arbitrary CDF $G(x)$, the CDF and PDF of the exp-G distribution with power parameter $a > 0$ are

$$\Pi_a(x) = G(x)^a \quad \text{and} \quad \pi_a(x) = a g(x) G(x)^{a-1},$$

respectively. This class of distributions is quite useful in several applications. In fact, Tahir and Nadarajah (2015) [1] cited more than seventy papers on exponentiated distributions in their Table 1.

First, the CDF of the MO-G distribution (1.2) admits the linear representation (Barreto-Souza *et al.*, 2013) [16]

$$(6.1) \quad F_{\text{MO}-\Gamma}(x) = \sum_{i=0}^{\infty} w_i^{\text{MO-G}} \Pi_{i+1}(x) = \sum_{i=0}^{\infty} w_i^{\text{MO-G}} G(x)^{i+1},$$

where the coefficients are (for $i = 0, 1, \dots$)

$$w_i^{\text{MO}-\Gamma} = w_i^{\text{MO}-\Gamma}(\theta) = \begin{cases} \frac{(-1)^i \theta}{(i+1)} \sum_{j=i}^{\infty} (j+1) \binom{j}{i} \bar{\theta}^j, & \theta \in (0, 1), \\ \theta^{-1} (1 - \theta^{-1})^i, & \theta > 1, \end{cases}$$

and $\bar{\theta} = 1 - \theta$.

Second, the linear combination for the Γ -G cumulative distribution (1.4) follows from Castellares and Lemonte (2015) [17] as

$$(6.2) \quad F_{\Gamma-G}(x) = \sum_{j=0}^{\infty} w_j^{\Gamma-G} \Pi_{a+j}(x).$$

Here,

$$w_j^{\Gamma-G} = w_j^{\Gamma-G}(a) = \frac{\varphi_j(a)}{(a+j)},$$

$$\varphi_0(a) = \frac{1}{\Gamma(a)}, \quad \varphi_j(a) = \frac{(a-1)}{\Gamma(a)} \psi_{j-1}(j+a-2), \quad j \geq 1,$$

and

$$\begin{aligned} \psi_{n-1}(x) = & \frac{(-1)^{n-1}}{(n+1)!} \left[H_n^{n-1} - \frac{x+2}{n+2} H_n^{n-2} + \frac{(x+2)(x+3)}{(n+2)(n+3)} H_n^{n-3} - \dots \right. \\ & \left. + (-1)^{n-1} \frac{(x+2)(x+3)\dots(x+n)}{(n+2)(n+3)\dots(2n)} H_n^0 \right] \end{aligned}$$

is the Stirling polynomial, $H_{n+1}^m = (2n+1-m)H_n^m + (n-m+1)H_n^{m-1}$ is a positive integer, $H_0^0 = 1$, $H_{n+1}^0 = 1 \times 3 \times 5 \times \dots \times (2n+1)$ and $H_{n+1}^n = 1$.

By inserting (6.2) in Equation (6.1) and via a result for a power series raised to a positive integer (Gradshteyn and Ryzhik, 2000) [18], the expansion for the cdf of the MOGG distribution reduces to

$$\begin{aligned} F_{\text{MO-}\Gamma\text{-G}}(x) &= \sum_{i=0}^{\infty} w_i^{\text{MO-}\Gamma} G(x)^{(i+1)a} \left[\sum_{j=0}^{\infty} w_j^{\Gamma-G} G(x)^j \right]^{i+1} \\ &= \sum_{i=0}^{\infty} w_i^{\text{MO-}\Gamma} G(x)^{(i+1)a} \sum_{j=0}^{\infty} c_{i+1,j} G(x)^j = \sum_{i,j=0}^{\infty} d_{i,j} \Pi_{(i+1)a+j}(x), \end{aligned}$$

where $d_{i,j} = d_{i,j}(a, \theta) = w_i^{\text{MO-G}} c_{i+1,j}(a)$, $c_{i+1,0}(a) = (w_0^{\Gamma-G})^{i+1}$ and, for $m \geq 1$, $c_{i+1,m}(a) = \frac{1}{mw_0^{\Gamma-G}} \sum_{r=1}^m [r(i+2) - m] w_r^{\Gamma-G} c_{i+1,m-r}(a)$.

By differentiating the last equation, the linear representation for the MOGG density holds

$$(6.3) \quad f_{\text{MO-}\Gamma\text{-G}}(x) = \sum_{i,j=0}^{\infty} d_{i,j} \pi_{(i+1)a+j}(x).$$

So, some structural properties of the proposed family can be determined from the double linear combination (6.3) and those properties of the exp-G distribution. In most applications, the indices i and j can vary up to a small number of terms.

6.2. Some quantities

Hereafter, let $T_{i,j} \sim \text{exp-G}[(i + 1)a + j]$. The n -th moment of X can be determined from (6.3) as

$$(6.4) \quad \mu'_n = E(X^n) = \sum_{i,j=0}^{\infty} d_{i,j} E(T_{i,j}) = \sum_{i,j=0}^{\infty} [(i + 1)a + j - 1] d_{i,j} \tau[n, (i + 1)a + j - 1],$$

where

$$\tau(n, a) = \int_{-\infty}^{\infty} x^n G(x)^a g(x) dx = \int_0^1 Q_G(u)^n u^a du.$$

Expressions for moments of several exponentiated distributions can be found in the papers cited in Tahir and Nadarajah (2015, Table 1). We give just one example from Equation (6.4) by taking the exponential distribution with rate $\lambda > 0$ for the baseline G. It follows easily as

$$\mu'_n = n! \lambda^n \sum_{i,j,m=0}^{\infty} \frac{(-1)^{n+m} [(i + 1)a + j] d_{i,j}}{(m + 1)^{n+1}} \binom{(i + 1)a + j - 1}{m}.$$

For empirical purposes, the shape of many distributions can be usefully described by the incomplete moments. These moments play an important role for measuring inequality. For example, the mean deviations and Lorenz and Bonferroni curves depend upon the first incomplete moment of the distribution. The n -th incomplete moment of X can be expressed as

$$(6.5) \quad m_n(y) = \int_{-\infty}^y x^n f_X(x) dx = \sum_{i,j=0}^{\infty} [(i + 1)a + j] d_{i,j} \int_0^{G(y)} Q_G(u)^n u^{(i+1)a+j-1} du.$$

The definite integral in (6.5) can be evaluated for most baseline G distributions.

The moment generating function (MGF) $M(t) = E(e^{tX})$ of X can be expressed from (6.3)

$$(6.6) \quad M(t) = \sum_{i,j=0}^{\infty} d_{i,j} M_{i,j}(t) = \sum_{i,j=0}^{\infty} [(i + 1)a + j] d_{i,j} \rho(t, (i + 1)a + j - 1),$$

where $M_{i,j}(t)$ is the MGF of $Y_{i,j}$ and

$$\rho(t, a) = \int_{-\infty}^{\infty} e^{tx} G(x)^a g(x) dx = \int_0^1 \exp\{t Q_G(u)\} u^a du.$$

The MGFs of several MOGa-G distributions can be determined from Equation (6.6). For example, the generating function of the MOGa-exponential with parameter λ (if $t < \lambda^{-1}$) is

$$M(t) = \sum_{i,j=0}^{\infty} [(i + 1)a + j] d_{i,j} B((i + 1)a + j, 1 - \lambda t).$$

7. CONCLUSIONS

A new family of distributions called the Marshall and Olkin-Gamma-G family with two shape parameters is introduced. The estimation of the unknown parameters is done via the maximum likelihood method and a simulation study is conducted to verify its adequacy. Additionally, the usefulness of the proposed family is shown empirically by means of two applications to real data. In fact, the new family can generate very competitive distributions with the same number of parameters than others constructed by existing classes.

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